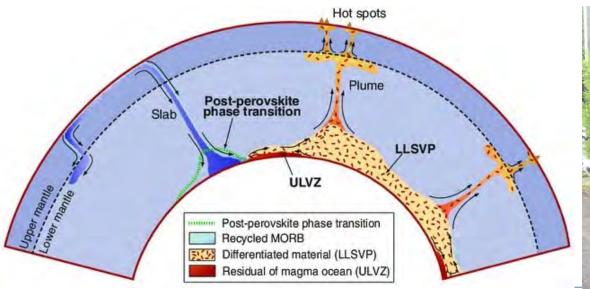
Mantle structures and their origins

Bernhard Steinberger

GFZ German Research Centre for Geosciences, Potsdam, Germany



GFZ

Helmholtz-Zentrum



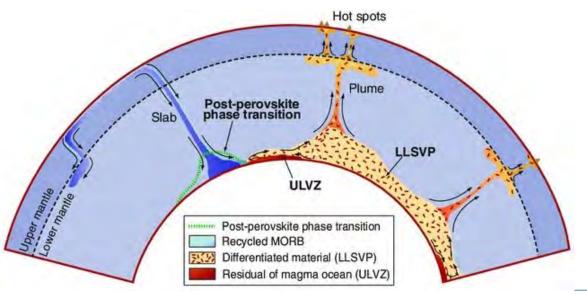
Figure from Deschamps, Li & Tackley, 2015

Sendai, Summer 2017 HELMHOLTZ

- Mantle contains thermal and chemical structure
- Convection creates thermal structure; erases chemical structure through mixing
- Melting separates materials and causes chemical structure
- Large uncertainties on mantle rheology

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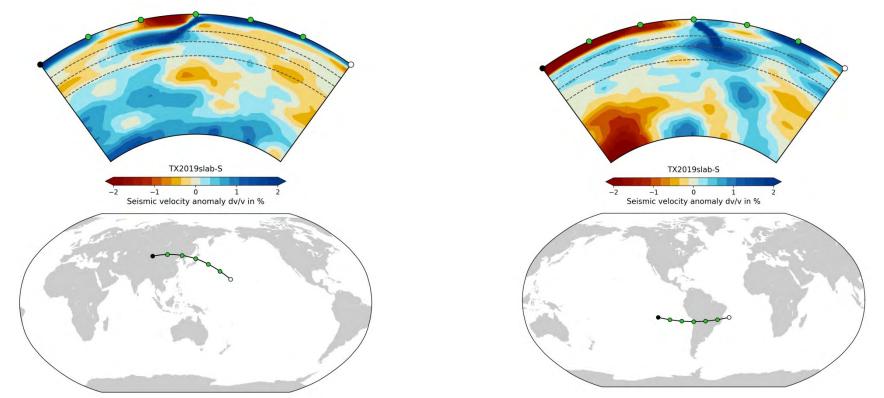
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Example for stagnating slab: Japan

Example for penetrating slab: South America



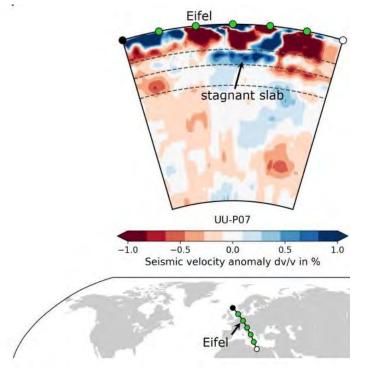
Cross sections from SubMachine (Hosseini et al.) https://www.earth.ox.ac.uk/~smachine

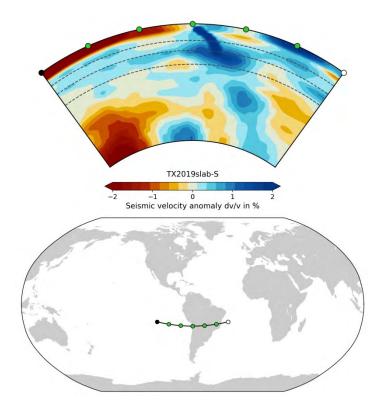


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Example for stagnating slab: Europe

Example for penetrating slab: South America

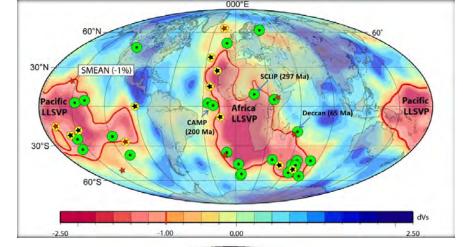


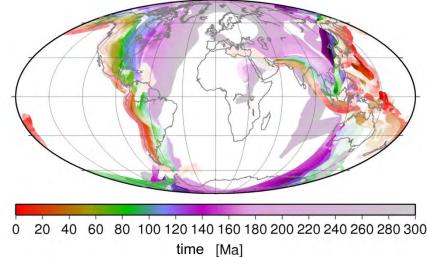






→ s-wavespeed anomalies in lowermost mantle, → reconstructed LIPs (green) → likely deep hotspots (stars) → after Torsvik et al. (2006)

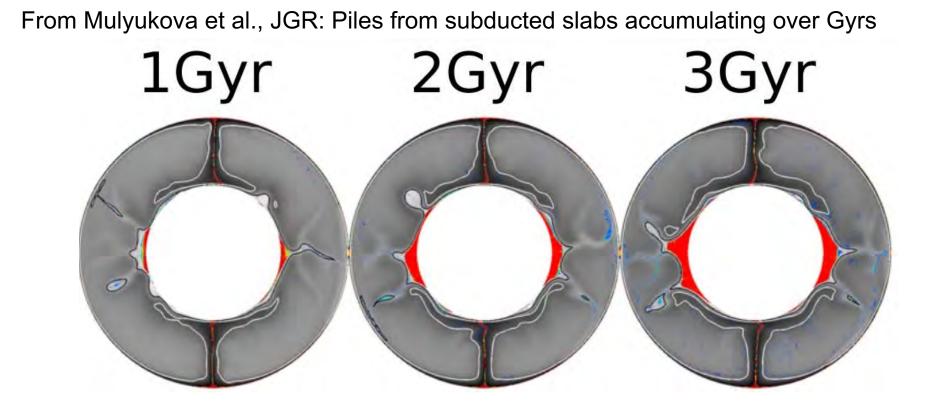




Subduction locations and amounts (color intensity) from a global plate reconstruction (see Steinberger and Torsvik, 2012, for details)

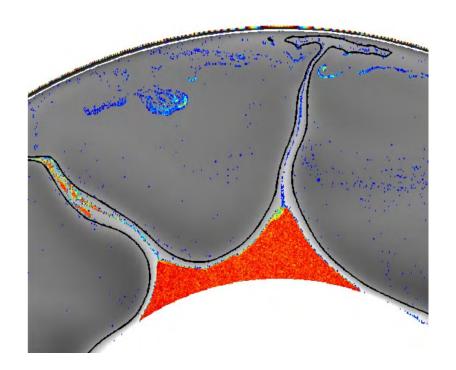


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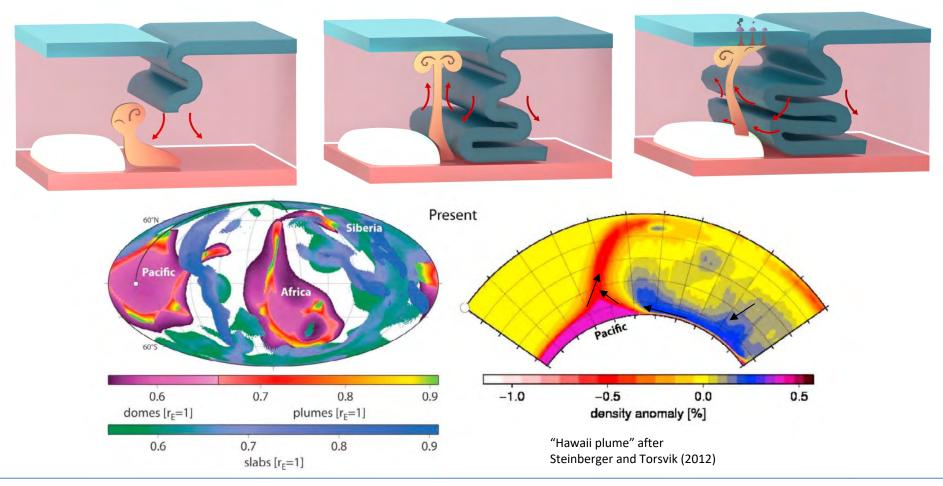




- Zoom-in on one of the piles
- plumes forming at cusps near margins
- plumes entraining pile materials



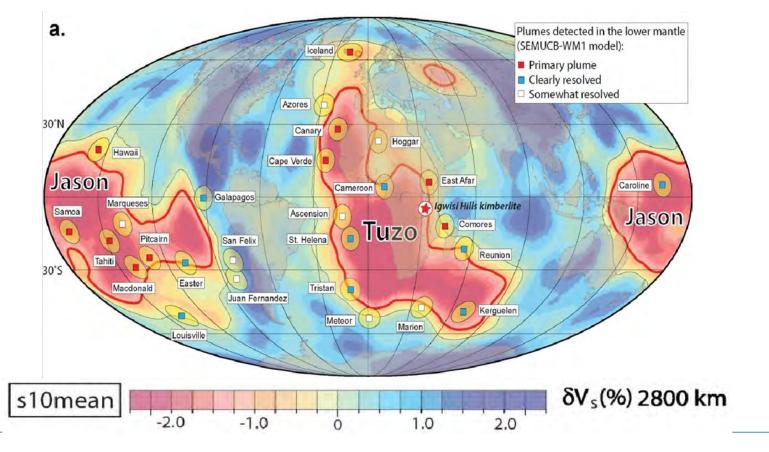




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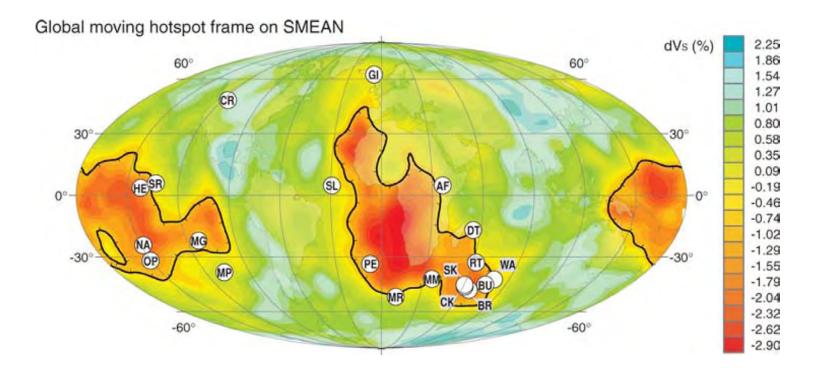
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Hotspots located near margins of Large Low Shear Velocity Provinces (LLSVPs) "Jason" and "Tuzo"





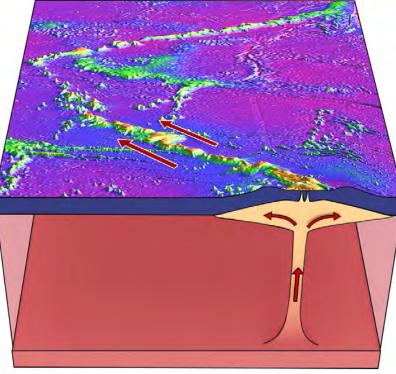
- Same is true also for reconstructed LIP locations
- Indicates spatial stability of LLSVPs back to at least 200 Myr, perhaps longer



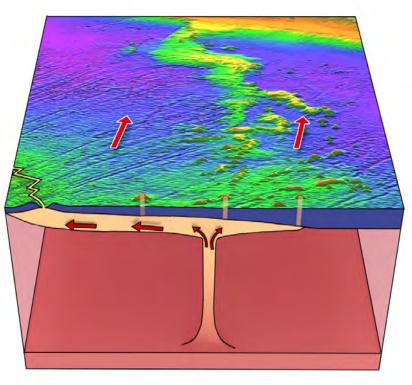




Plume-Lithosphere interaction – intraplate: Hotspot track and swell:



- Interaction with ridge (Gassmöller et al., G-Cubed, 2016)









TX2019slab-S (Lu et al., 2019)

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-1.0

2000

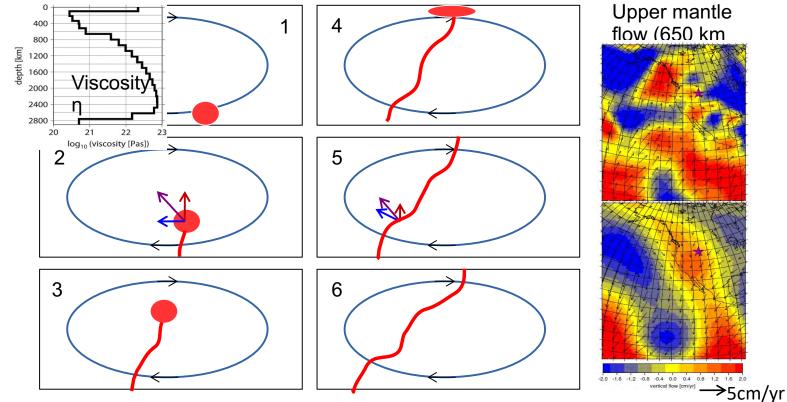
1000

topography [m]

P.0

0.5

0.0 anomaly [%] 5.0-

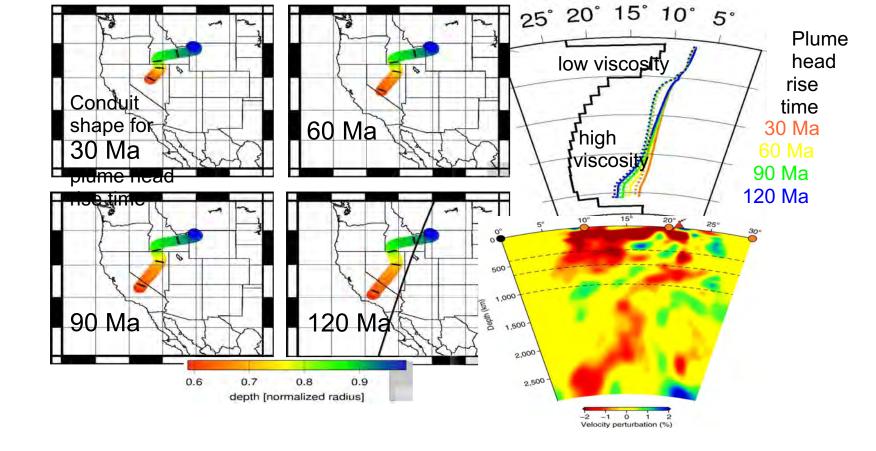


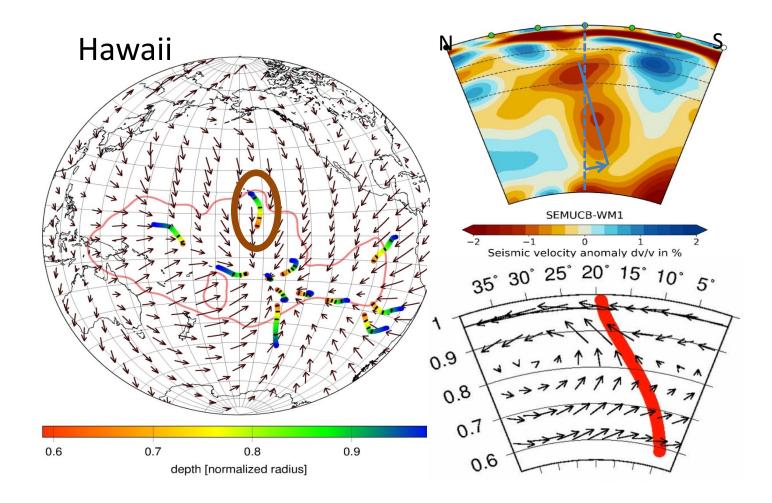
Plume head rises ($v_{rise} \sim 1/\eta$) gets advected by large-scale flow (density anomalies from tomography) How fast?

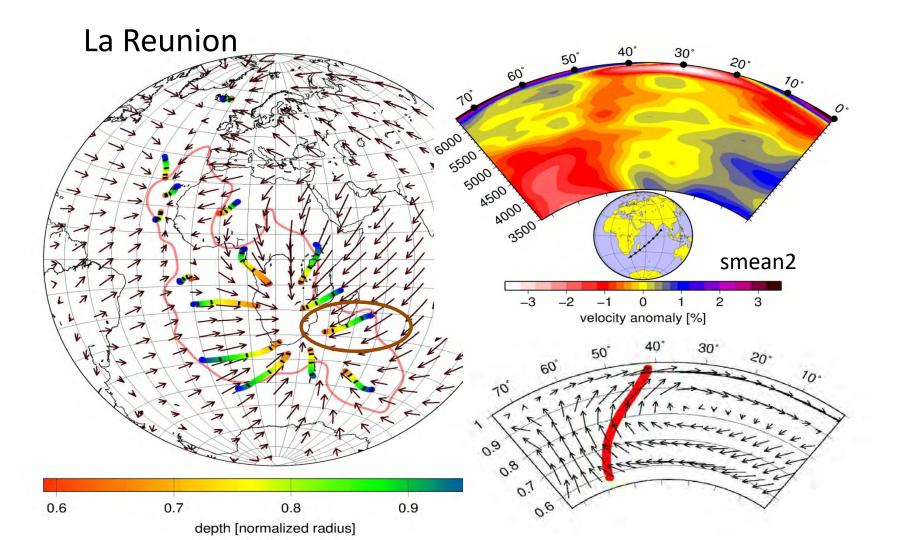
 \rightarrow vary total rise time (prescribed)

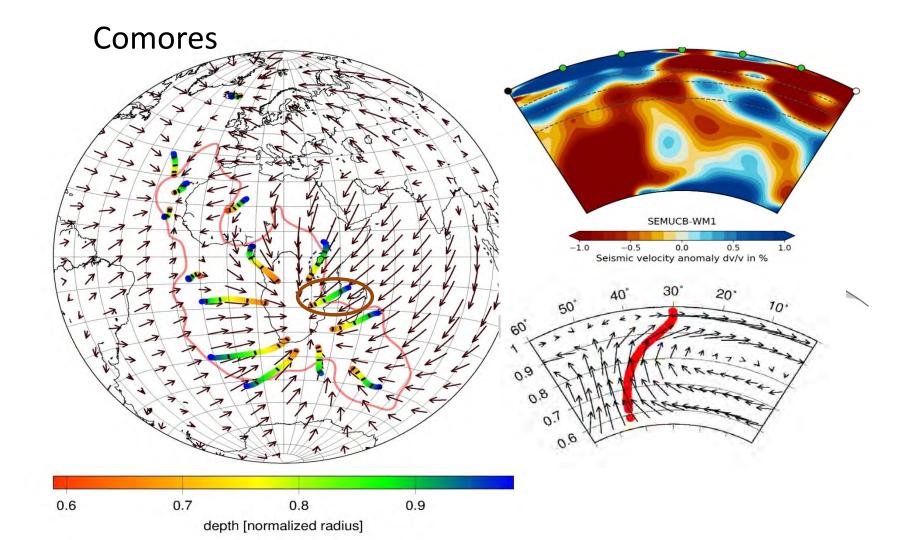
Plume conduit also rises (but less fast) + gets advected by large-scale flow

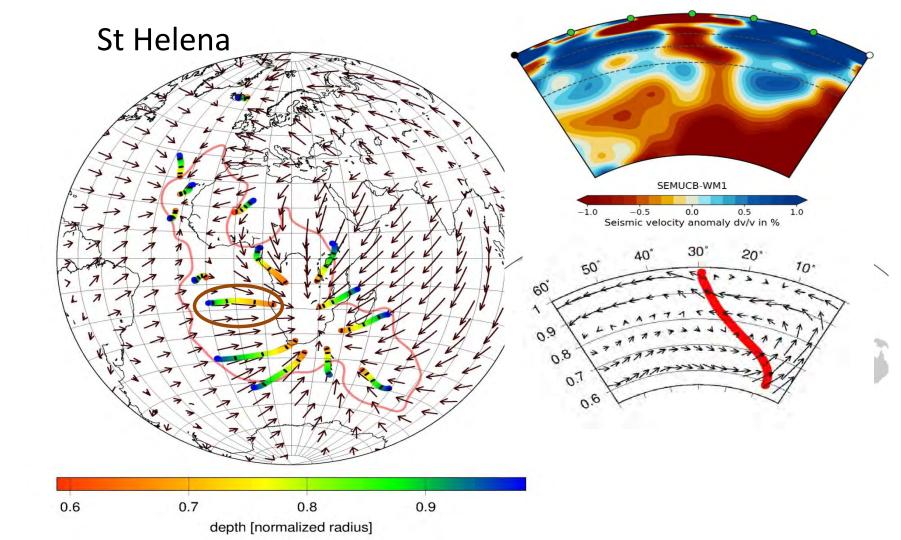
Lower mantle flow (2600 km depth)

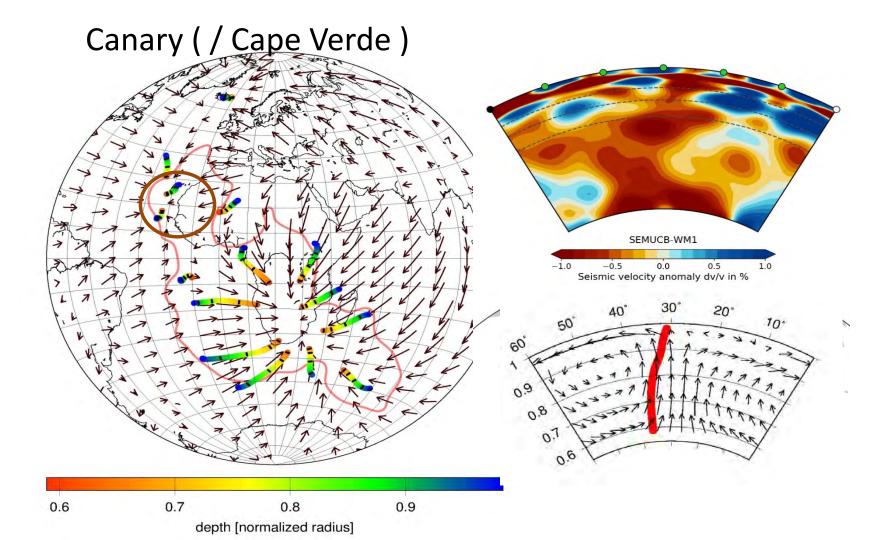






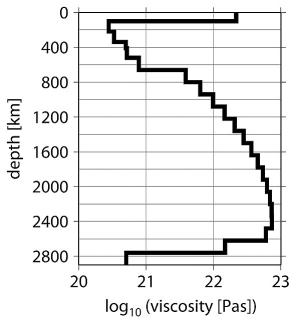






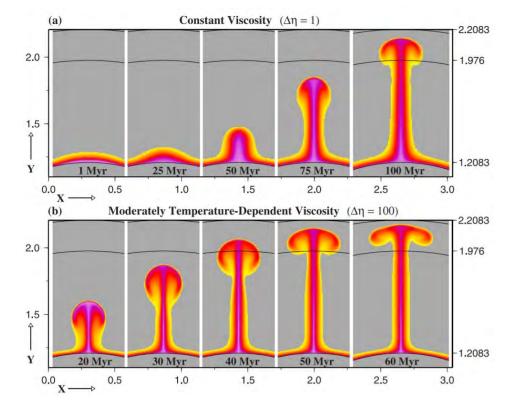
Radial viscosity:

- Postglacial rebound
- Geoid modelling
- Slow motion of hotspots



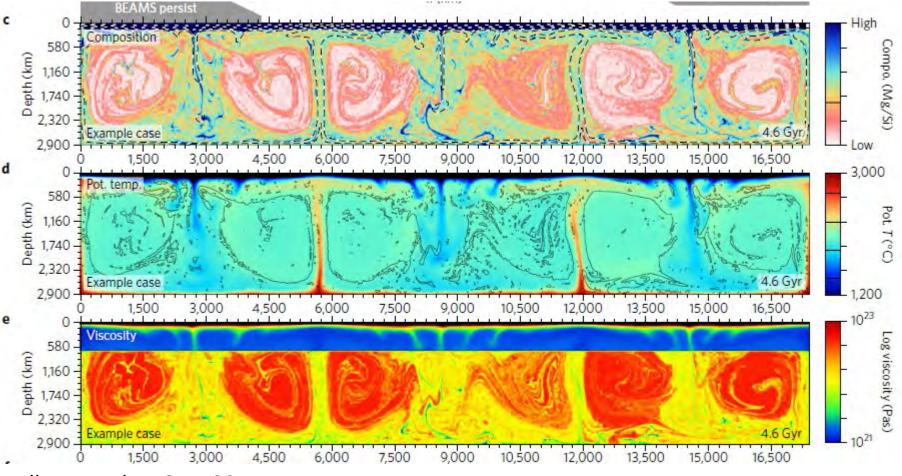
Steinberger & Calderwood (2006)

Temperature dependent viscosity Causing narrow plume conduits



Lin and van Keken (2006)

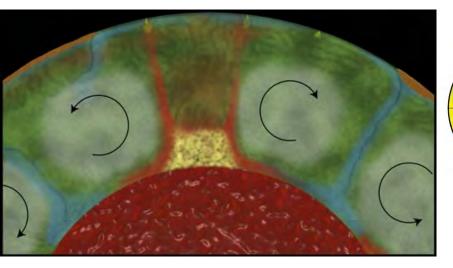
Composition dependent viscosity: bridgmanite-enriched ancient mantle structures



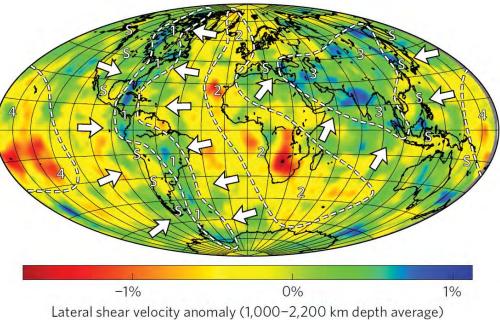
Ballmer et al., NGeo, 2017

Cartoon view of BEAMS

Suggested distribution: Between dashed lines, below arrows



Ballmer et al., NGeo, 2017



Integrated strain dependent viscosity LBF=Load-bearing framework

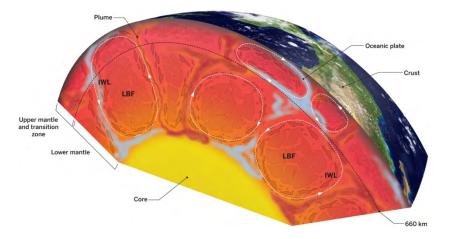


ILLUSTRATION: C. BICKEL/SCIENCE Illustrating Girard et al., Science, 2016

- Where Bridgmanite is interconnected, mantle is stronger.
- However, strong shear may lead to interconnection of weaker Magnesiowustite





Summary

- Thermal structure created through convection: slabs going down, plumes coming up.
- Chemical structure: Crust generated through melting.
 Accumulated in piles on base of mantle? Entrained in plumes?
- Ancient (primordial) structure in piles? In BEAMS?
- Complicated rheology: Viscosity depth dependent, temperaturecomposition- strain(rate), strain history dependent?

HELMHO

